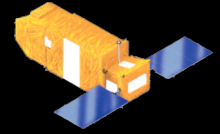
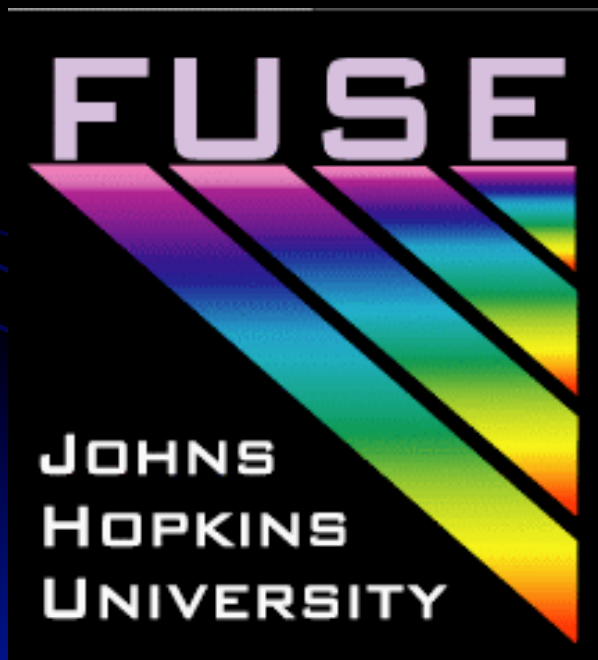


# Far Ultraviolet Spectroscopic Explorer



## The FUSE Observatory: Status and Update of One Wheel Operations

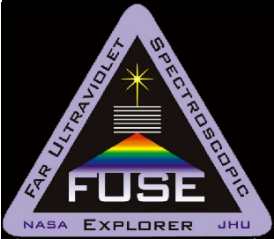


**Bill Blair**

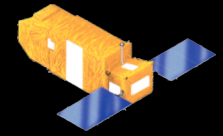
**FUSE Deputy-PI and**

**Chief of Observatory Operations**

**FOAC Meeting, March 31, 2006**



# FUSE--A Brief History



FUSE-Dec. 1999



FUSE-Feb. 2002

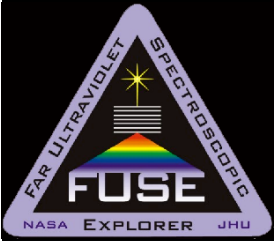


~~FUSE-June 2005~~

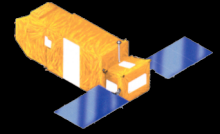
**October 2005**



FUSE-Mar. 2004



# FUSE--A Brief History



FUSE-Dec. 1999



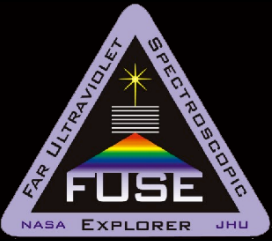
FUSE-Feb. 2002



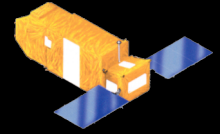
FUSE-Mar. 2004



January 2006!

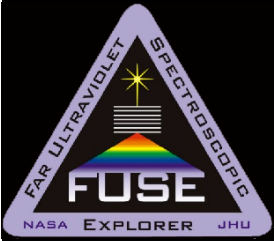


# Outline/Themes



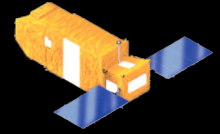
**FUSE at KSC, May 1999.**

- Background/Terminology
- Status/Improvements since last FOAC meeting
- Momentum Management
- Operations/MP Improvements
- Performance Indicators
- Long Observation Strategy
- ACS/IDS Code Updates
- Possible Future Improvements
- Questions/Issues
- CaFUSE/MAST Status (Dixon)

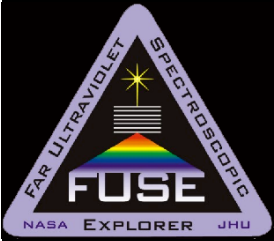


# One-Wheel Ops

## A Primer

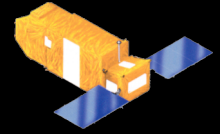


- Attitude Control System (ACS) is the S/C software that controls pointing.
- Only Wheel remaining is the Skew Reaction Wheel.
  - +/- 6500 rpm top speed (+/-21 Nms).
  - Higher wheel speeds mean more gyroscopic torques when slewing.
  - We plan so as to keep this below +/-15 Nms.
    - But Reality does not always follow what is planned!
- Three Magnetic Torquer Bars (MTBs) mounted on the body axes of the satellite, need to share duty between control and momentum unloading for the wheel.
- Two-axis Magnetometers (TAMs) provide attitude knowledge to +/-2 degrees.
- Fine Error Sensor (controlled by the Instrument Data System computer) provides Fine Pointing Data (FPDs) to the ACS.

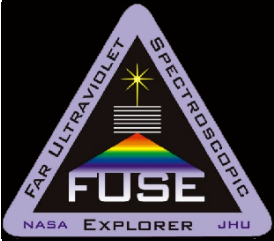


# Mission Status/Overview

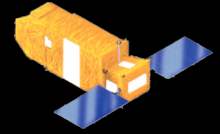
## (Since last FOAC meeting)



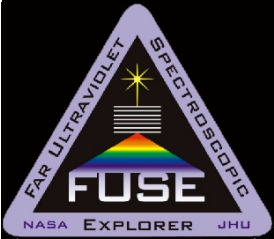
- Science Ops officially restarted Nov. 1, 2005.
- FES-B performance continues to be nominal.
- UPRM Ground station fixed and performance has been steady.
  - Increasing frequency for preventive maintenance.
- Automated/streamlined Safemode recovery slew calculation process. (Next pages.)
- Nov-Dec 2005 “Demonstration period” was difficult, but successful.
  - Got started on some science as we continued to learn.
  - Developed concept of (and tools for) “momentum interventions.”
  - Ops team pulled out all the stops, but nearly burned out.
  - Had to “park” the satellite over the holidays.



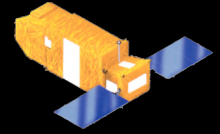
# Mission Status/Overview (continued)



- January AAS meeting was a big success.
  - Special Oral session; 43 papers total.
- Operations have improved dramatically so far in 2006.
  - Fairly subtle change in MP software (mid-Dec) made a BIG difference.
  - Other improvements to MP software are having an effect.
  - JHU press release generated touting “Return to Operations.”
- NASA Senior Review was submitted Mar. 20.
  - Strong case for continued exciting science from FUSE.
- Latest and greatest ACS code (E33) and accompanying IDS script rewrites developed and tested; load occurred earlier this week.
- Instituting a revised methodology for aligning the channels.
- Victoria Proceedings (ASP) have just been delivered to the printer.
- CalFUSE/Reprocessing work has made good progress. (See end.)



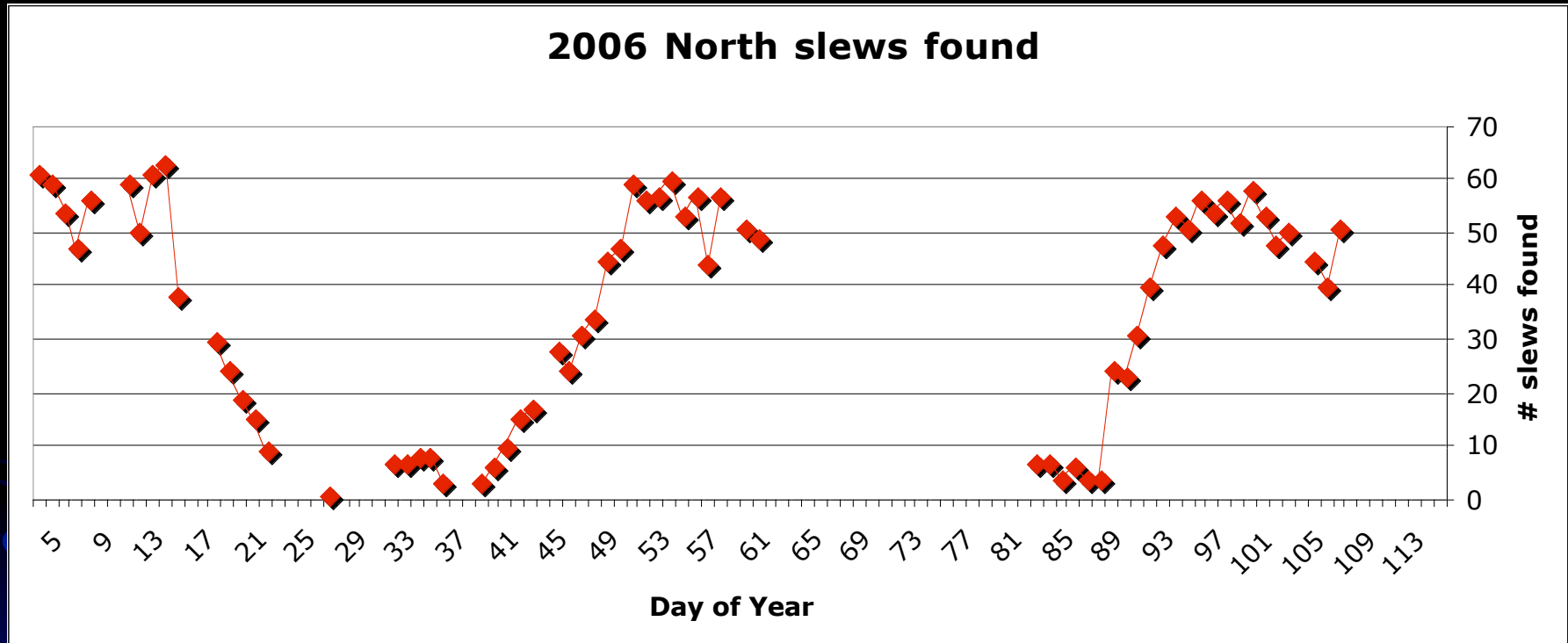
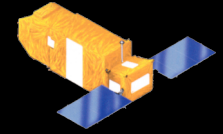
# LVLH Safe Mode



- LVLH (Local Vertical Local Horizontal) is a nadir-pointing, non-inertial safe mode.
- Because it is not an inertial pointing mode, transitions back from LVLH to an inertial pole-pointing (pick up point) can be difficult to find.
  - Nominally “safe” slews are found with the HDS.
  - Number of opportunities “per day” are quite variable with time.
  - TDRS or other contact times must be arranged to monitor slew progress and attempt intervention if needed.
- Once at an orbit pole, we must “match momentum” with a planned timeline before picking up.
  - Typically end recovery slew with moderately high momentum on the wheel.
- Process has been automated since last FOAC, so potential slews are available if/when needed. (Substantial effort.)
- New (more robust and predictable) procedure just tried this week; may supplant previous method eventually.



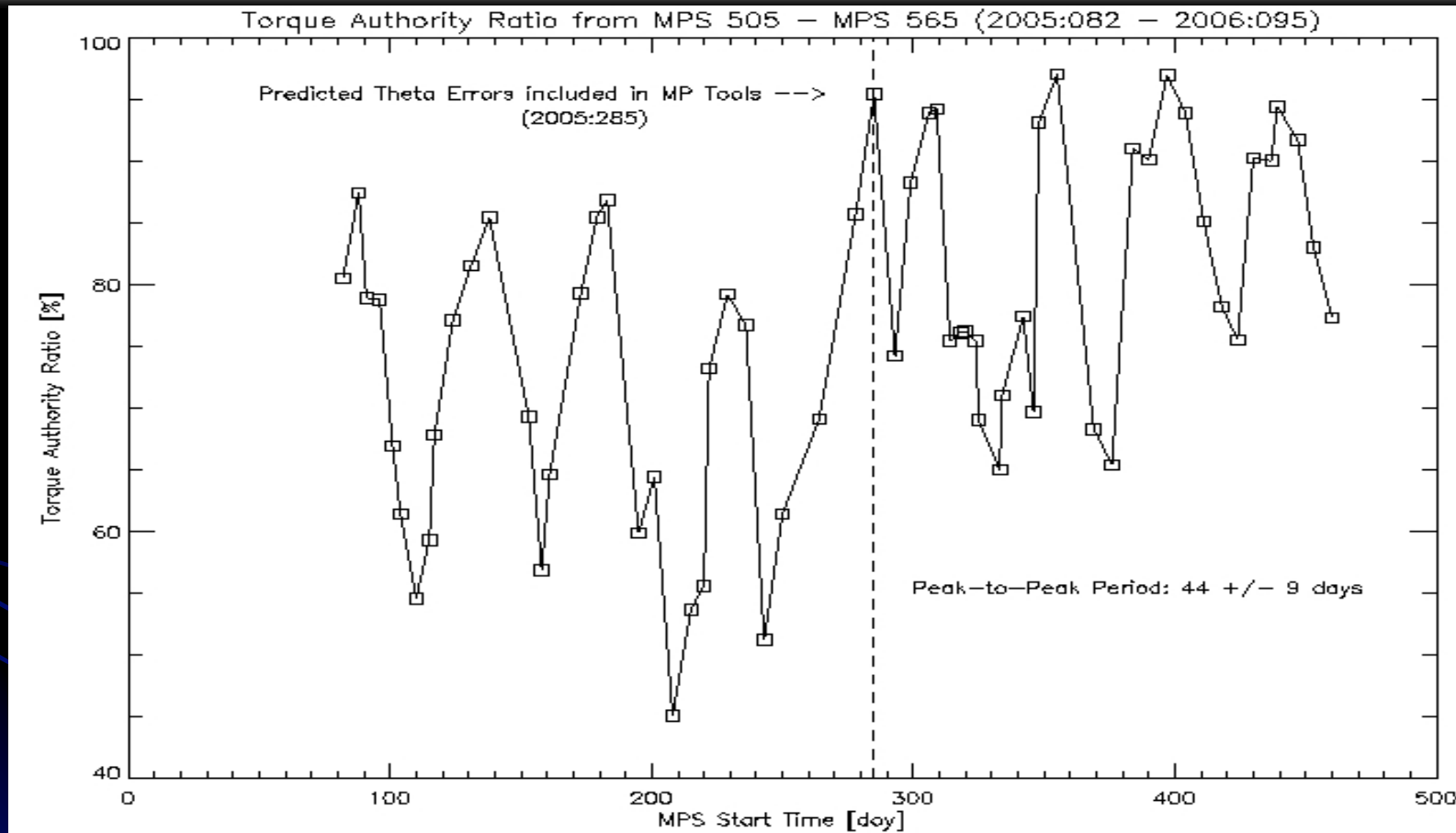
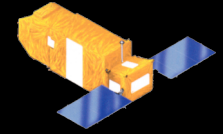
# LVLH Recovery Slews vs. Time



**Empirical result: Number of recovery slew available per day varies with ~45 day period.**

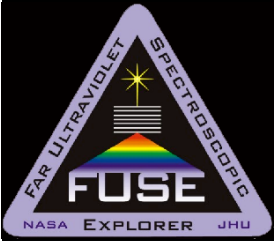
**(M. Kochte)**

# MPS TA vs. time

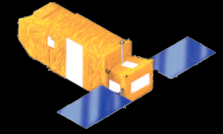


- ✓ Roughly Periodic oscillation between relatively stable and less stable conditions.
- ✓ Note general improvement in Oct. 2005 with the advent of improved planning tools.

(H. Calvani)

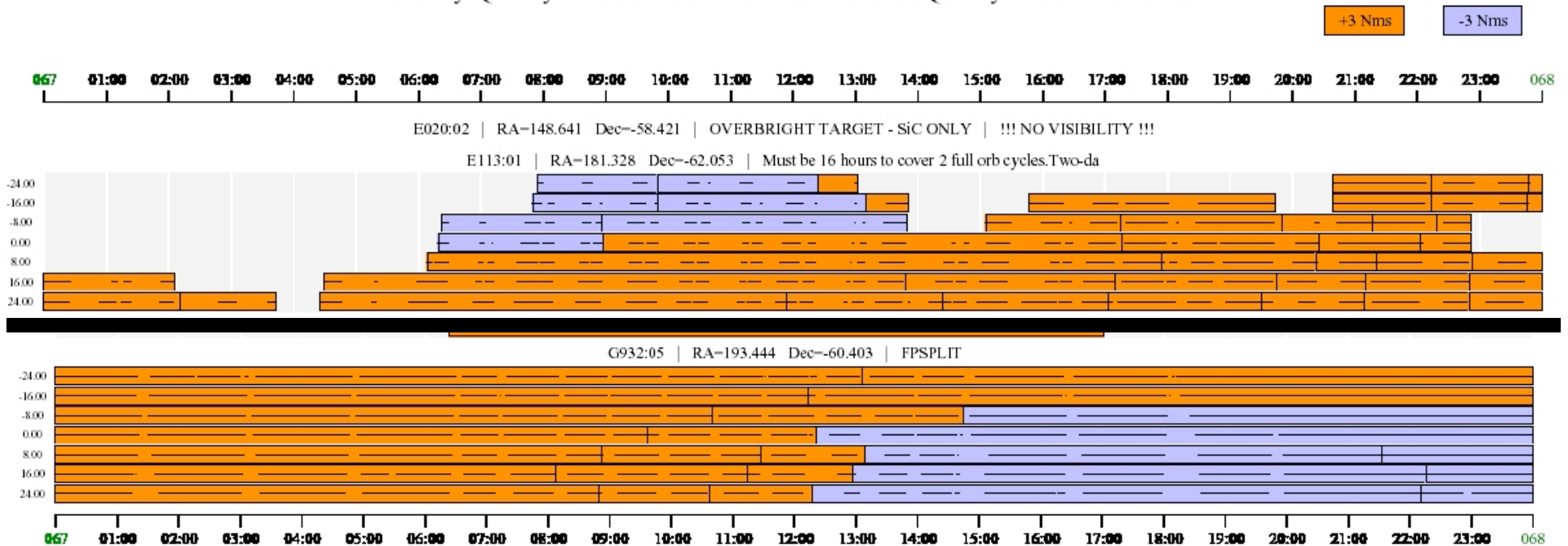


# MP Tool Improvement

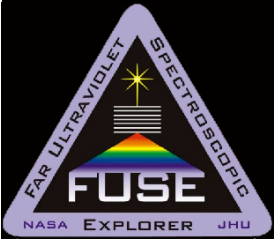


Visibility Timeline from 2006:067:00:00:00 to 2006:068:00:00:00

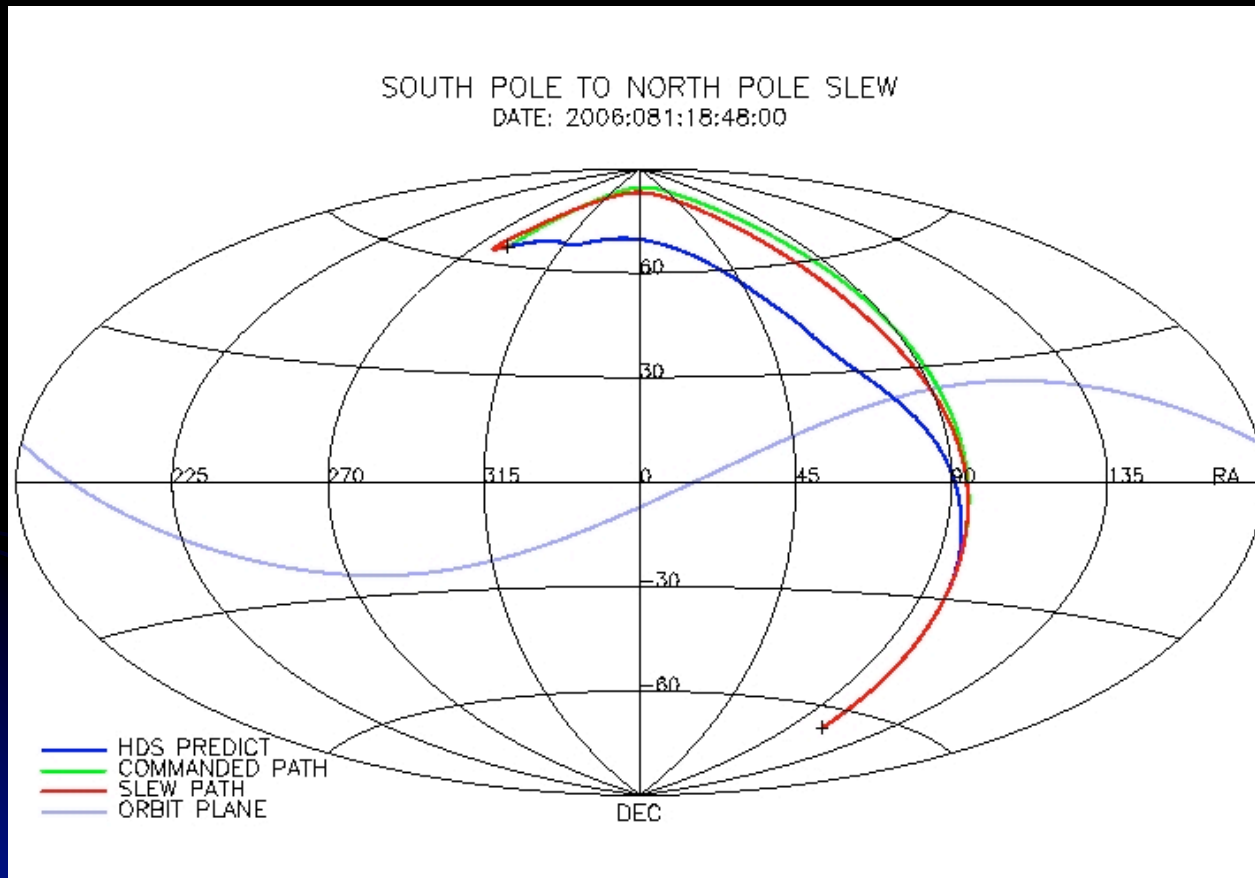
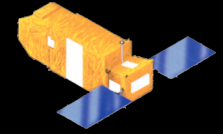
Stability Quality Factor = 100.00% Momentum Quality Factor = 78.95%



- Assimilating info on stability, momentum rate of change, and flexibility available from roll angle variations, on a target-by target basis and as a function of time, provides planners with nearly all relevant data at once.



# Hemisphere Crossing Slews



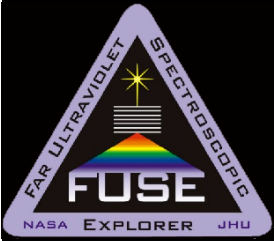
**Timing and matching of starting momentum is critical to success.**

**Also need to manage momentum to rejoin the timeline in the new hemisphere.**

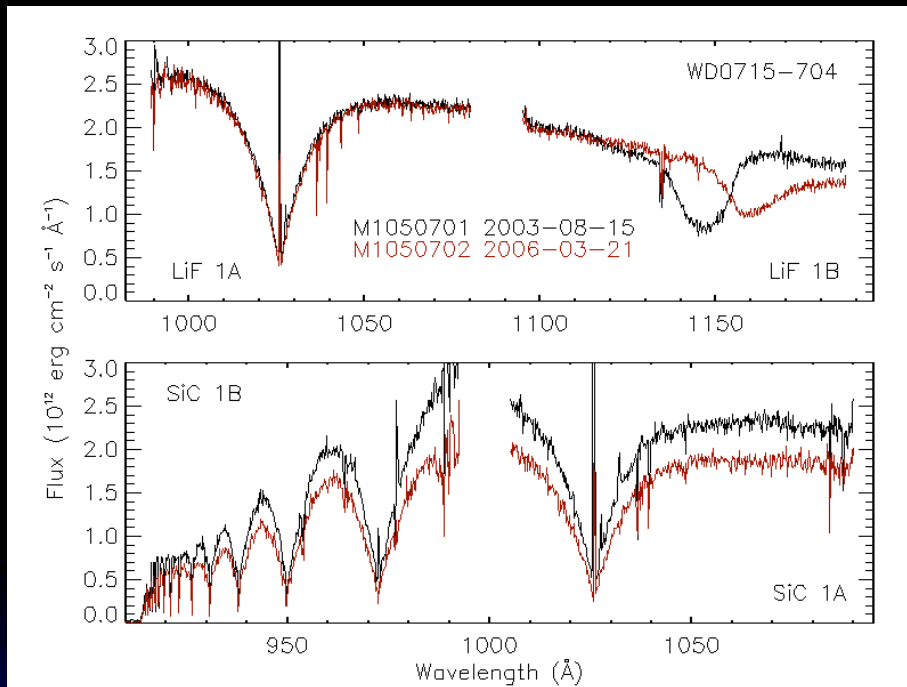
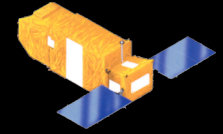
**Two successful slews and one failure to date. Also bailed on two others (went to LVLH and recovered to other pole instead of attempting slew) because of stability concerns.**

- Preparation for these slews is time consuming and matching starting momentum is tricky. >> Don't do them frivolously!

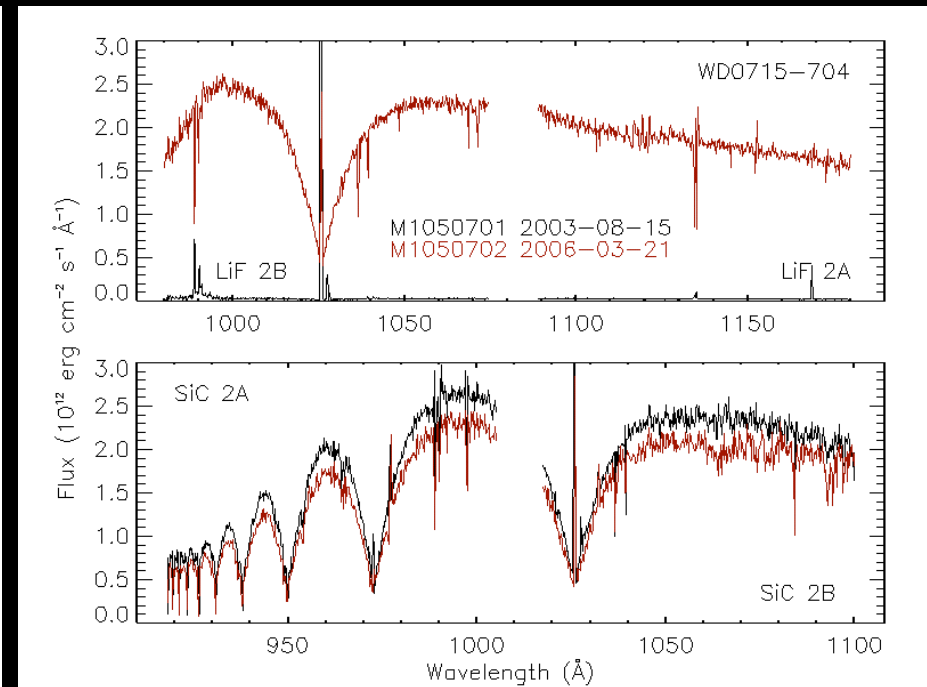
**(T. Ake)**



# Science Instrument Status



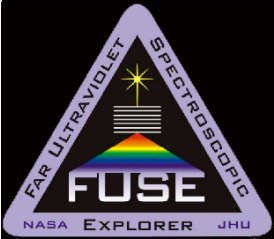
## Detector 1



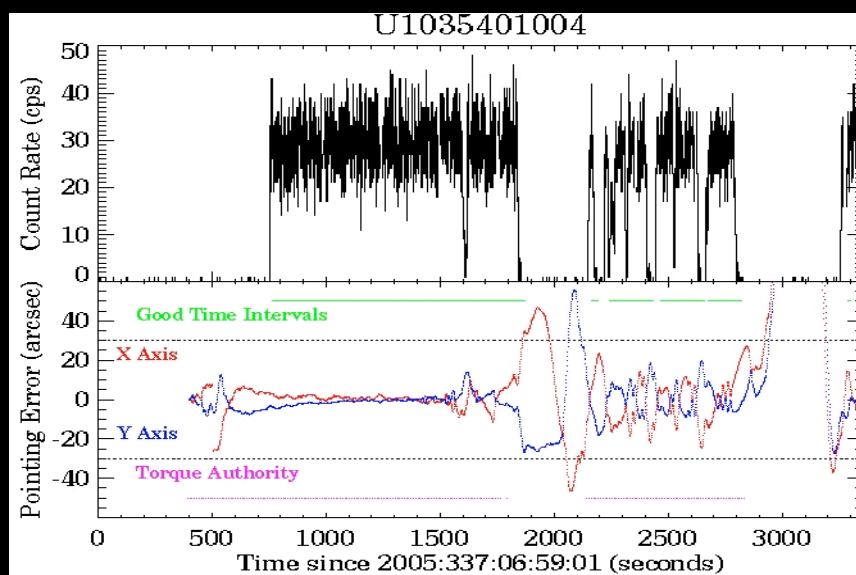
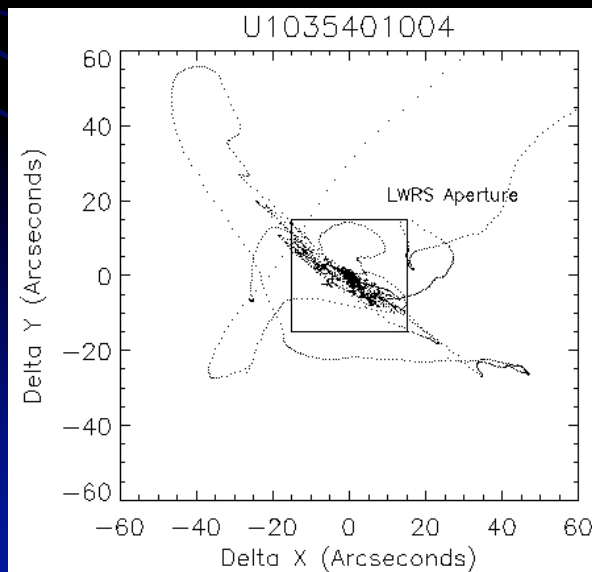
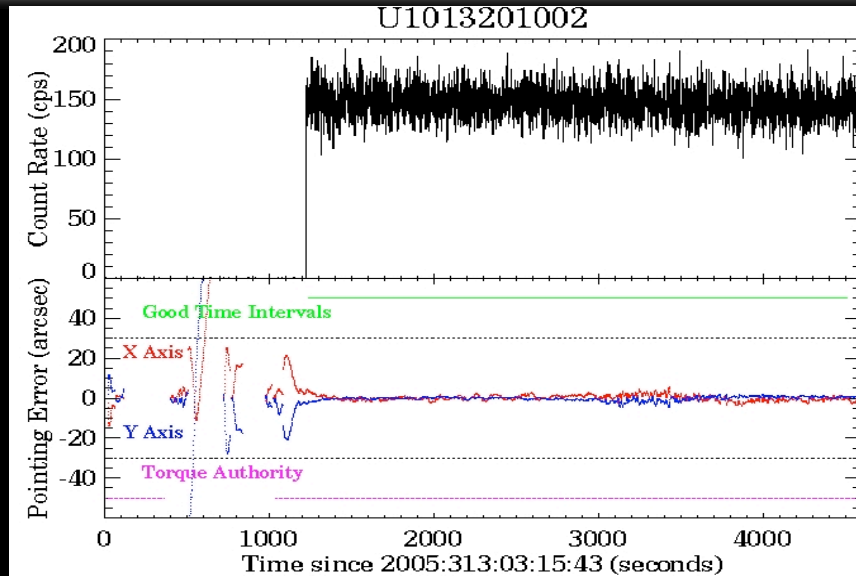
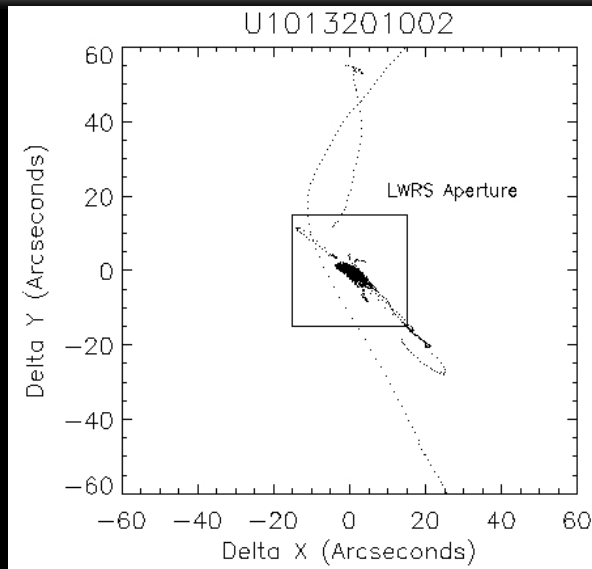
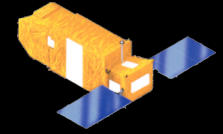
## Detector 2

- M10507, WD0715-704, comparison of data taken in Aug. 2003 and Mar. 2006. (~2.6 years apart)
- LiF1-no change; SiC1 ~20%; SiC2 ~10% decreases.
- Consistent with what has been seen earlier in the mission.

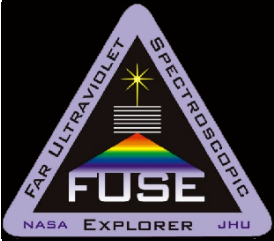
(V. Dixon)



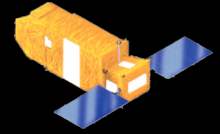
# System Performance



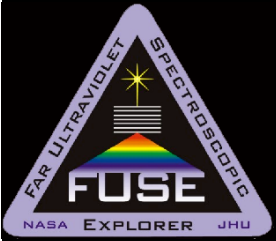
(T. Ake)



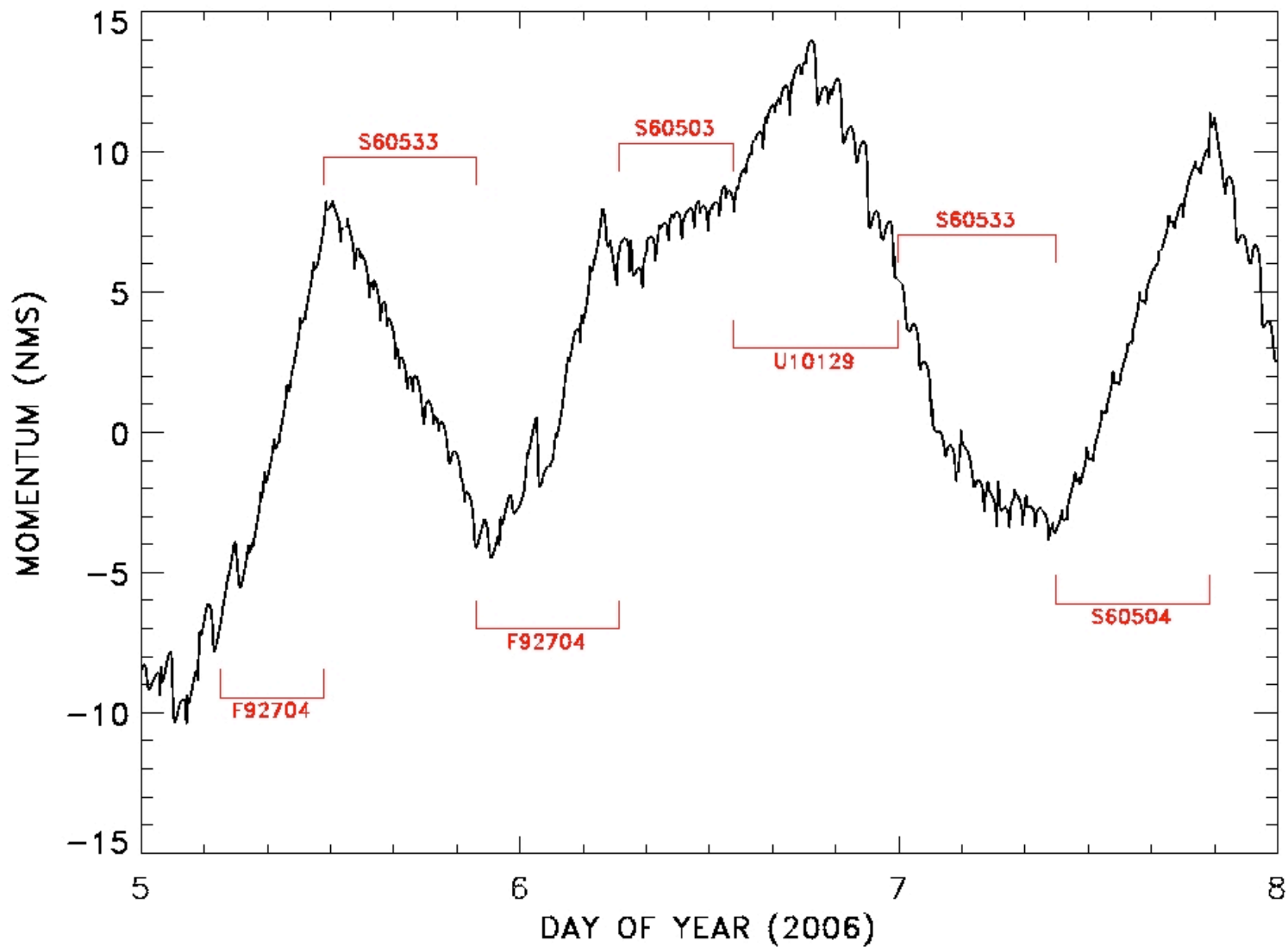
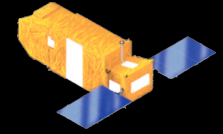
# Momentum Management



- With one wheel, careful management of momentum is critical to operations.
- Momentum is managed primarily by selection of pointing direction as a function of time.
- Predictive tools, such as the Hybrid Dynamic Simulator, are only accurate to a point. There are non-deterministic aspects that drive simulation and reality apart.
- There are many parameters within the control system that can be tweaked to improve performance, but some of them interact in ways that make accurate prediction of performance difficult.

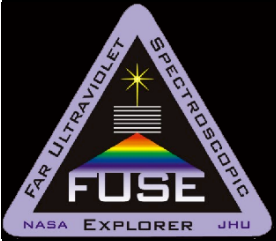


# Momentum Management Using Target Positions

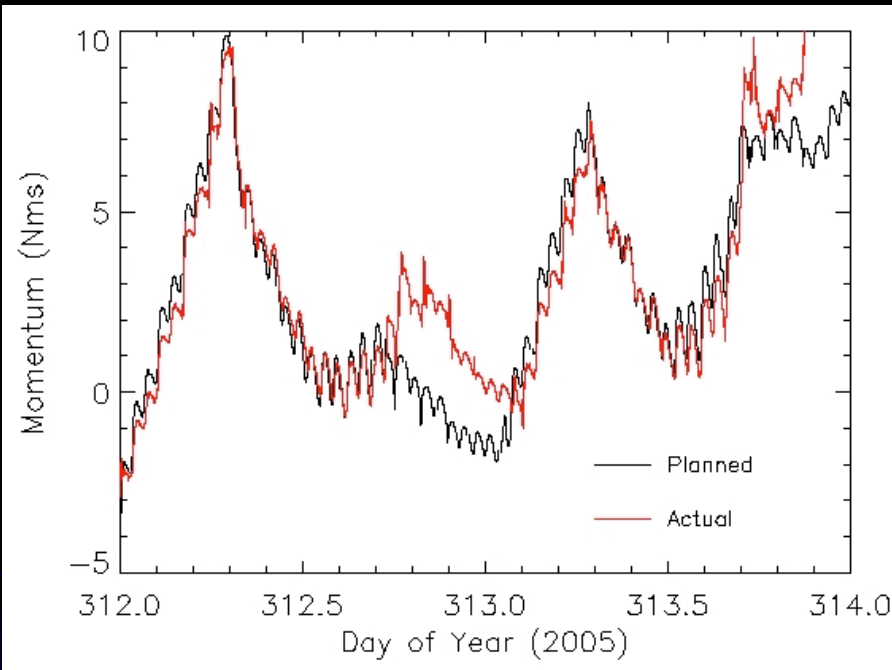
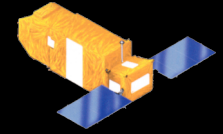


(T. Ake)

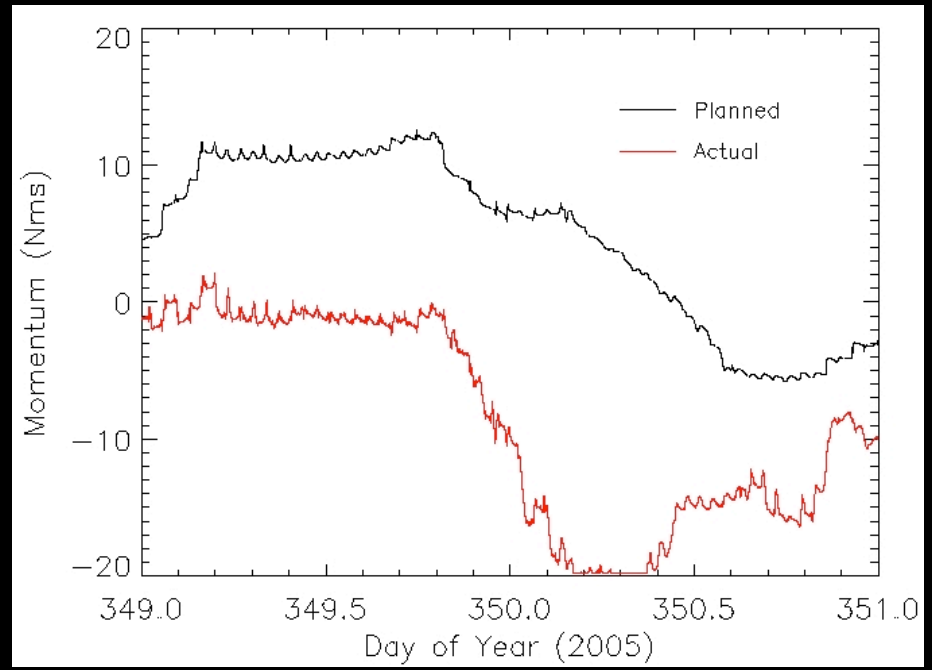




# Momentum Mgmt. Challenges



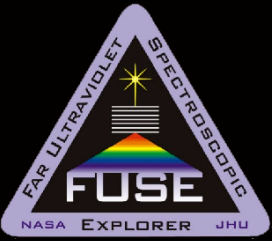
**Relatively Good Behavior  
Period**



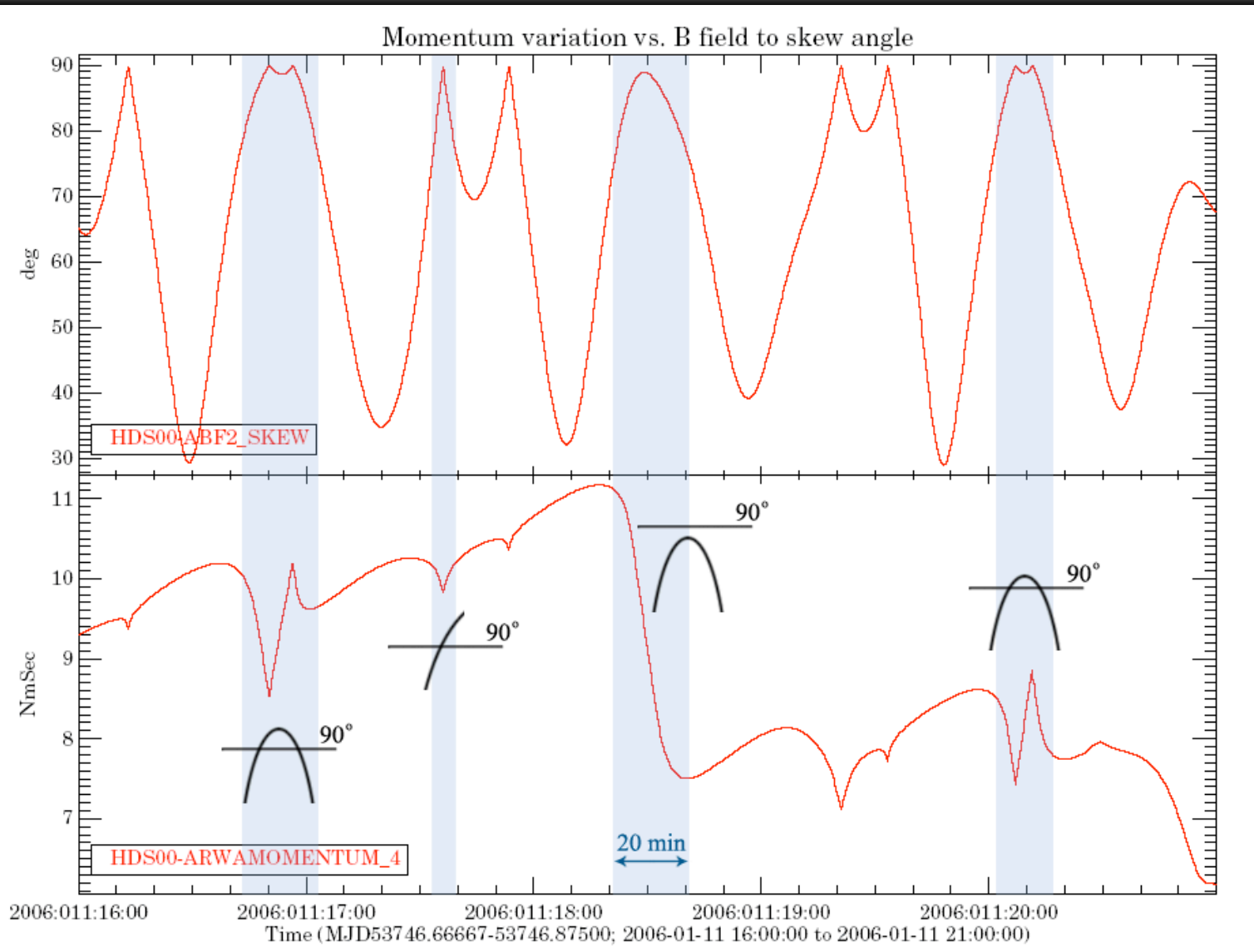
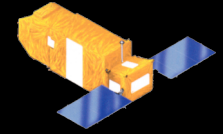
**Poorly Behaved Period.**

**A timely “momentum  
intervention” might have  
improved this situation.**

**(T. Ake)**



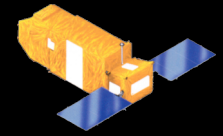
# B-field-to-Skew angle Sensitivity



(T. Civeit)



# January 2006 Performance



## FUSE Performance Summary – data from telemetry and MPS's

Total time available (ksec)	2,251.61
Total time scheduled (ksec)	1,140.23
% time sched	<b>50.64%</b>
Total time guiding (ksec)	676.51
% time guiding vs. req	76.11%
% time guidi	<b>59.33%</b>

Hemisphere crossing slews		
Date	Direction	Success?
1/18	N to S	Y
1/31	S to N	N

Spacecraft Downtime		
Type	# occurrences	Time (ks)
Planned LVLH	1	302.4
Unplanned LVLH	0	0.0
Mom mgmt	5	206.0
Other	0	0.0
<b>TOTALS</b>	<b>6</b>	<b>508.4</b>

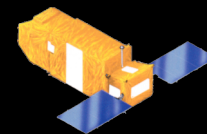
Program Category	Scheduled Time (ksec)	Guiding Time (ksec)	# obs sched
Prime	309.0	182.9	13
Legacy	16.0	12.0	1
Survey	141.6	86.8	9
Calibration	5.0	3.9	1
U programs	284.3	214.3	17
S programs	384.3	176.5	20
<b>TOTALS</b>	<b>1,140.2</b>	<b>676.5</b>	<b>61</b>

**LVLH = “Local Vertical Local Horizontal” (Nadir-pointing Safe Mode).**

**Analysis courtesy of Alice Berman, FUSE MP.**



# February 2006 Performance



## FUSE Performance Statistics – data from telemetry and MPS's

Total time in February (ksec)	<b>2,332.8</b>
Time avail for sched (ksec)	<b>2,075.5</b>
Actual time scheduled (ksec)	<b>1,222.9</b>
% time scheduled in all Feb	<b>52.4%</b>
% time sched of avail time	<b>58.9%</b>
Actual time guiding	<b>452.9</b>
% time guiding in all February	<b>19.4%</b>
% time guiding vs scheduled	<b>37.0%</b>
% time guiding vs available	<b>21.8%</b>

Hemisphere crossing slews		
Date	Direction	Success?
None	N/A	N/A

Spacecraft Downtime		
Type	# occurrences	Time (ks)
Planned LVLH	1	257.3
Unplanned LVLH	2	318.4
Mom mgmt	6	203.8
ACS Hold	2	70.1
<b>TOTALS</b>	<b>11</b>	<b>849.6</b>

2/24  
2/1, 2/11  
2/6, 2/8, 2/21, 2/22, 2/23  
2/2, 2/16

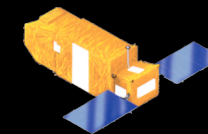
Program Category	Scheduled Time (ksec)	Guiding Time (ksec)	# obs sched	% guiding time
Prime	150.2	51.0	10	11.3%
Legacy	48.6	12.4	5	2.7%
Survey	446.7	194.2	22	42.9%
Calibration	24.9	0.9	6	0.2%
U programs	278.5	113.9	17	25.2%
S programs	274.1	80.5	12	17.8%
<b>TOTALS</b>	<b>1,222.9</b>	<b>452.9</b>	<b>72</b>	<b>100.0%</b>

Successful Observations	# obs
Successful	32
Partial success	2
Missed	22
Failed	16
<b>TOTAL</b>	<b>72</b>

Analysis courtesy of Alice Berman, FUSE MP.



# March 2006 Performance (March 1-22 only)



## FUSE Performance Statistics – data from telemetry and MPS's

Total time in March (ksec)	<b>1,987.2</b>
Time avail for sched (ksec)	<b>1,987.2</b>
<i>Actual time scheduled (ksec)</i>	<b>934.3</b>
% time scheduled in all March	<b>47.0%</b>
% time sched of avail time	<b>47.0%</b>
<i>Actual time guiding</i>	<b>658.8</b>
% time guiding in all March	<b>33.2%</b>
% time guiding vs scheduled	<b>70.5%</b>
<i>% time guiding vs available</i>	<b>33.2%</b>

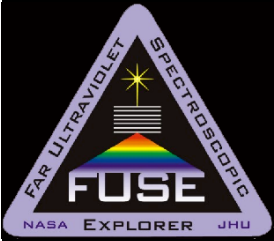
Hemisphere crossing slews		
Date	Direction	Success?
3/22	S->N	S

Spacecraft Downtime		
Type	# occurrences	Time (ks)
Planned LVLH	0	0.0
Mom mgmt	4	138.3
E33 load	0	0.0
ACS Hold	1	5.2
Unplanned LVLH	1	228.3
<b>TOTALS</b>	<b>6</b>	<b>371.8</b>

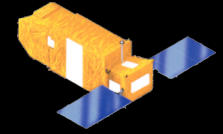
Program Category	Scheduled Time (ksec)	Guiding Time (ksec)	# obs sched	% guiding time
Prime	381.7	261.5	27	39.7%
Legacy	0.0	0.0	0	0.0%
Survey	308.2	246.1	16	37.4%
Calibration	45.6	11.7	3	1.8%
U programs	95.5	60.6	8	9.2%
S programs	103.4	78.8	8	12.0%
<b>TOTALS</b>	<b>934.3</b>	<b>658.8</b>	<b>62</b>	<b>100.0%</b>

Successful Observations	# obs
Successful	43
Partial success	0
Missed	9
Failed	10
<b>TOTAL</b>	<b>62</b>

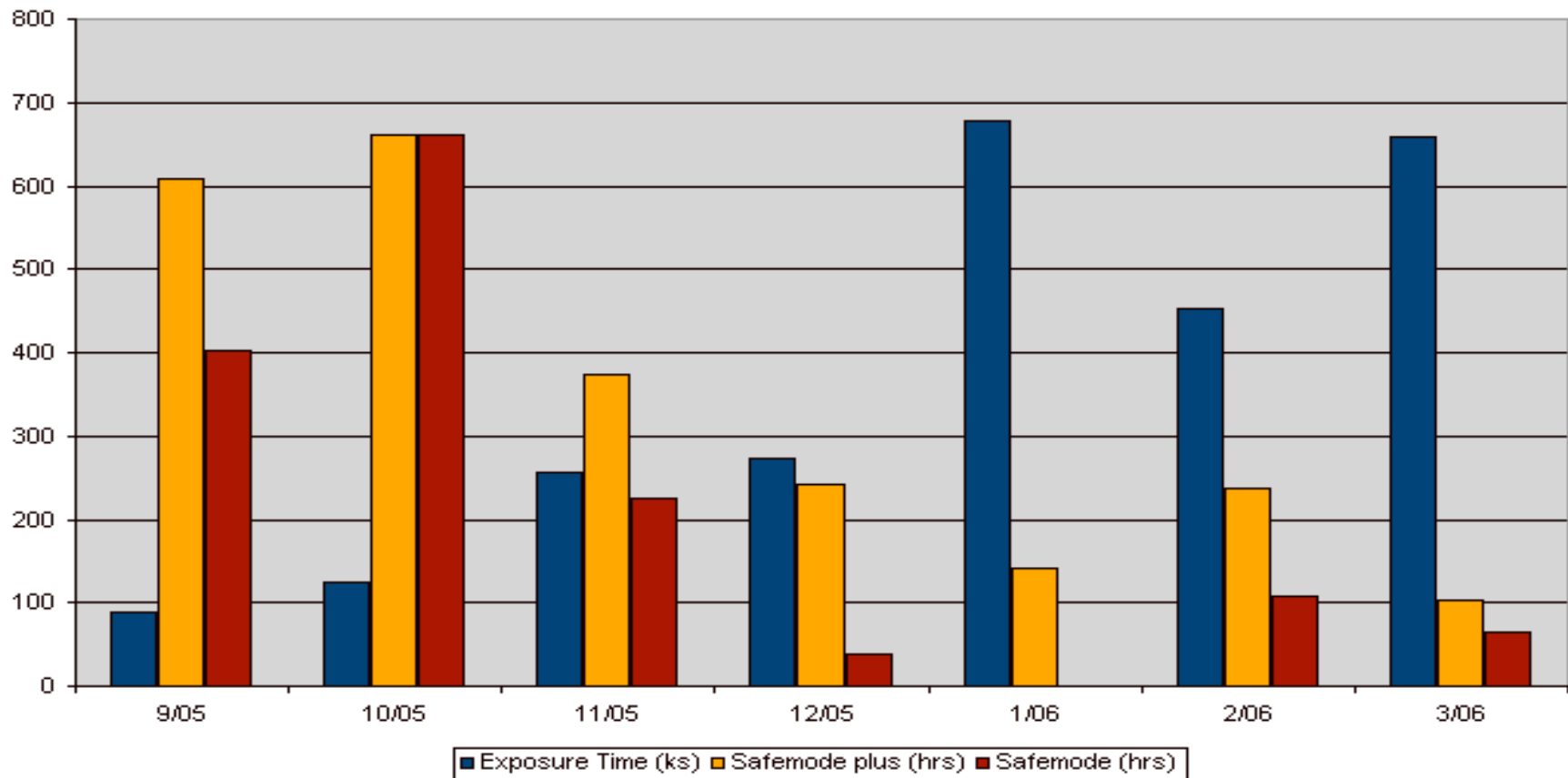
Analysis courtesy of Alice Berman, FUSE MP.



# Operations Improvements

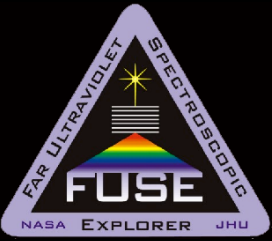


FUSE Performance Metrics  
(March numbers 3/1-3/22 only)

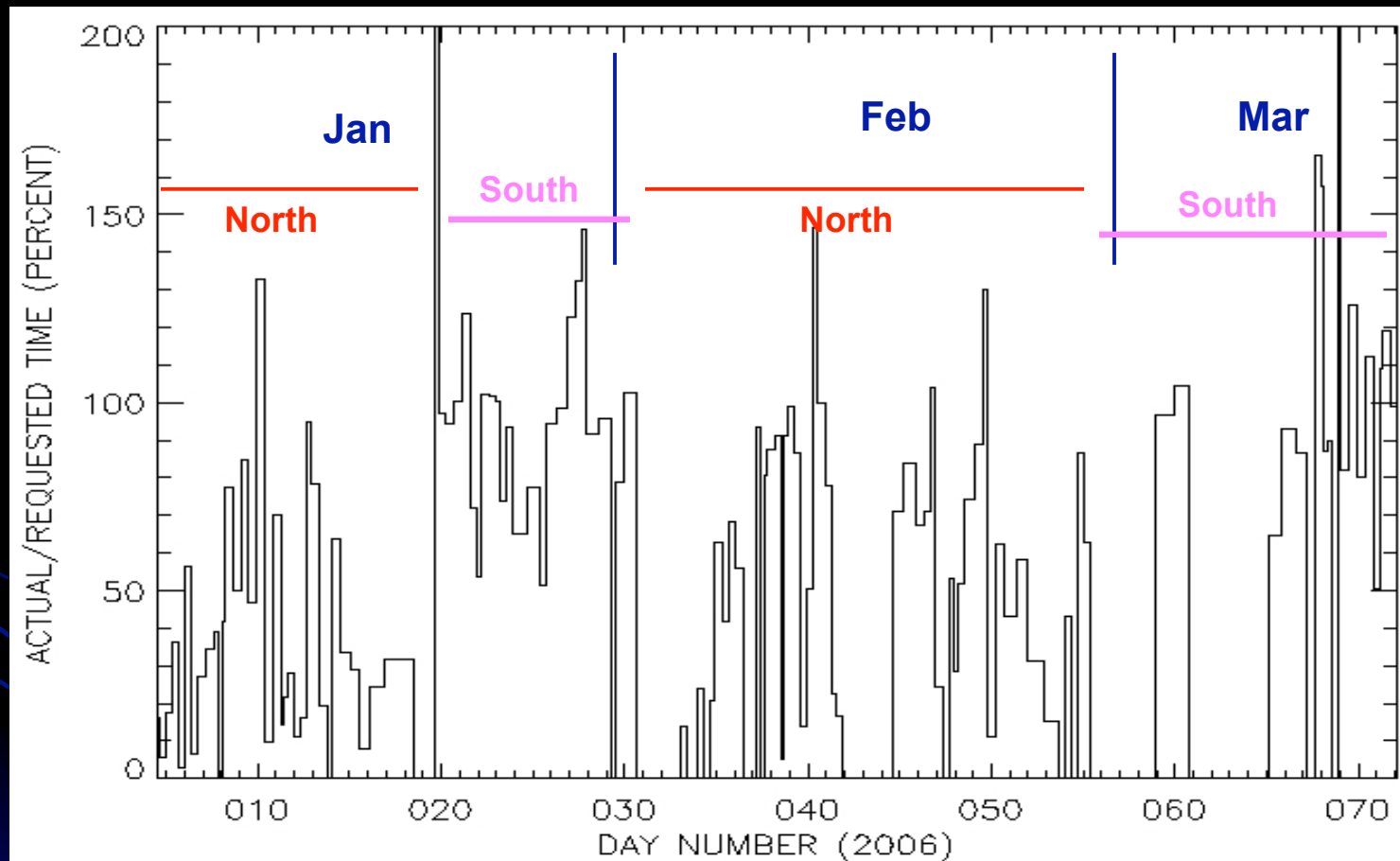
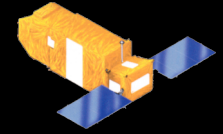


Yellow: Momentum mgmt and “planned” safemode time.

(A. Berman)

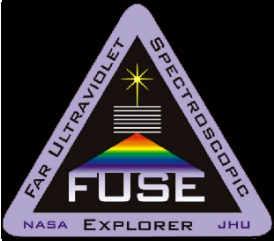


# Jan-mid Mar 2006 Actual vs Requested Times

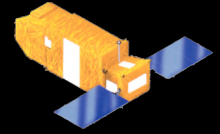


- By scheduling more time than requested, we typically get close to the requested times even with acquisition and guiding glitches.

(T. Ake)

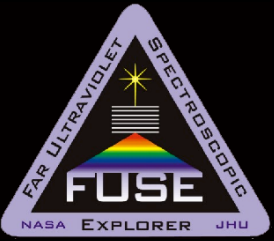


# Long Observations

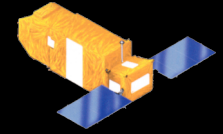


- Substantial effort invested in trying to scope this out but still not able to make very quantitative statements.
- There are target and scheduling condition combinations that allow long (upwards of 100 ks) observations to be scheduled in one period (say over a few days or so).
- Many targets have a distribution of visibility windows with 1-3 larger windows of ~50 ks and a number of smaller windows.
  - Requires a mechanism to lock down certain windows for certain targets.
  - The more of these there are, the larger the probability of conflicts.
  - Could easily drive otherwise unnecessary N-S changes.
- Clearly some dependence on declination ( $>60^\circ$  better than  $50-60^\circ$ ).
- In SR06, we said “~10 obs.  $> 100$  ks, with several as high as 200 ks” could be done per year.
- Bottom line: it depends on the specific targets, and on the priority assigned to them. (How hard should we work to get them?)

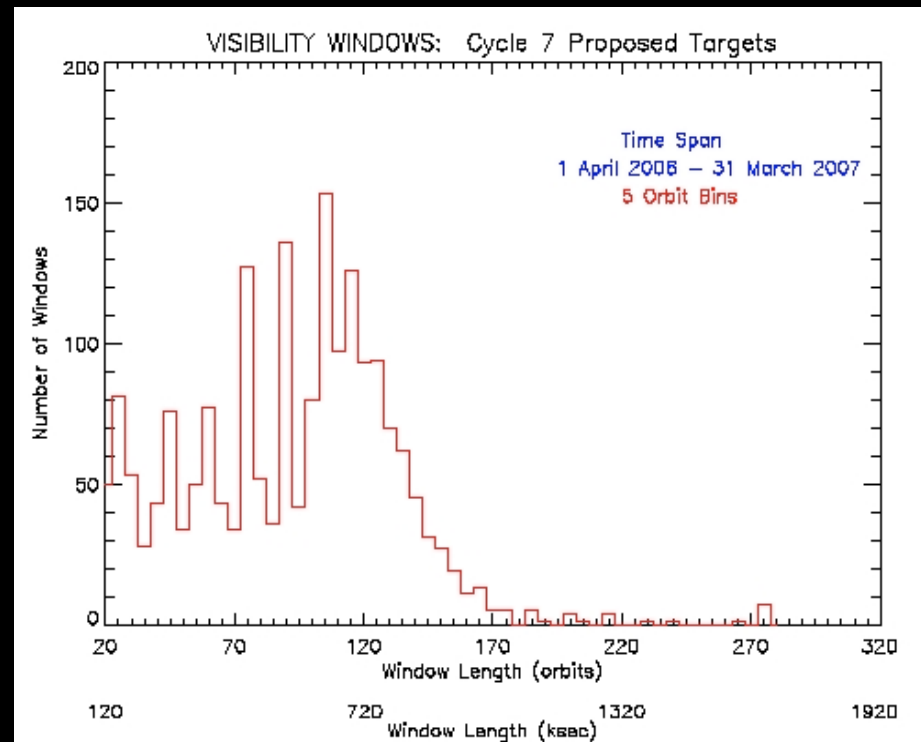
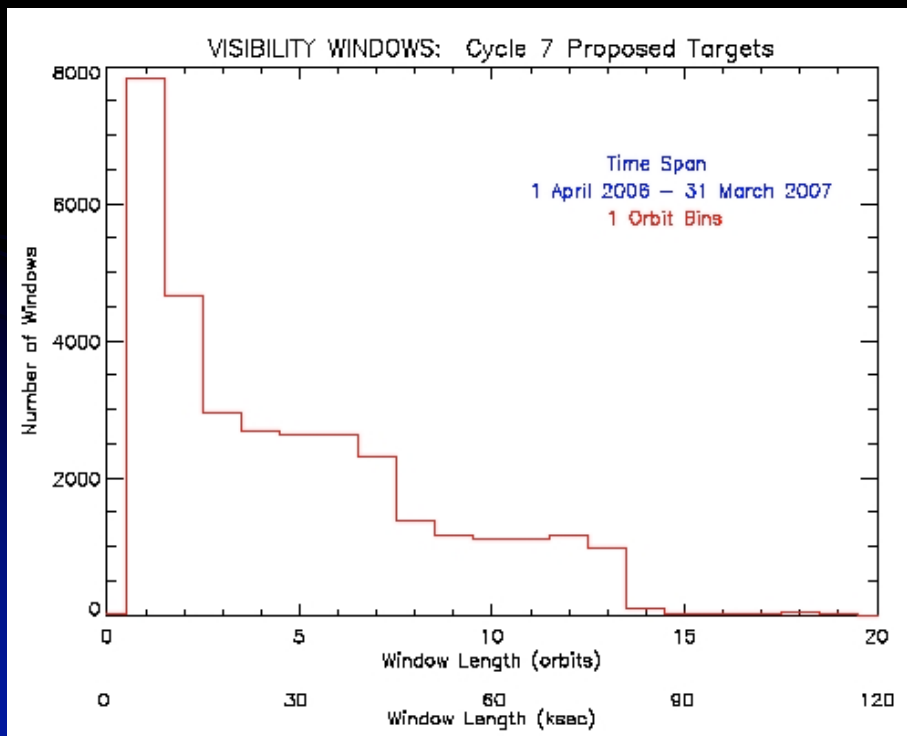




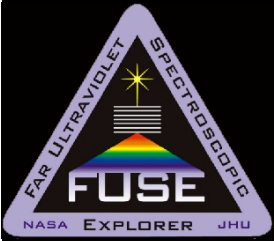
# Long Observations: Visibility Window Stats



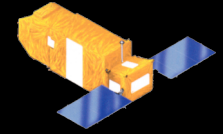
- Using all proposed Cy 7 targets (as representative).
- Stability periods only--No Momentum mgmt included.



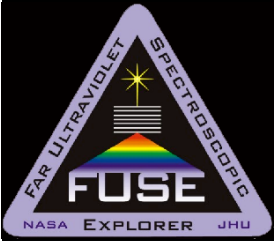
(M. England)



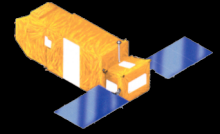
# ACS E33 / IDS v3.02



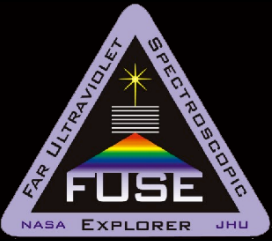
- New ACS flight s/w and new IDS s/w and scripts were loaded to FUSE earlier this week.
- Much improved attitude information sharing between ACS and IDS.
  - Fewer instances of bad FPD data corrupting ACS controller and causing loss of attitude.
- Revised Torque Distribution Algorithm at *B-to-S* angles near 90 degrees.
  - Fewer large and/or unexpected jumps in momentum compared with predicted behavior >> fewer momentum interventions needed.
- Medium gain controller--better slew and acquisition behavior.
- Improved/simplified LVLH recovery slews.
  - Not dependent on cyclical behavior with time.



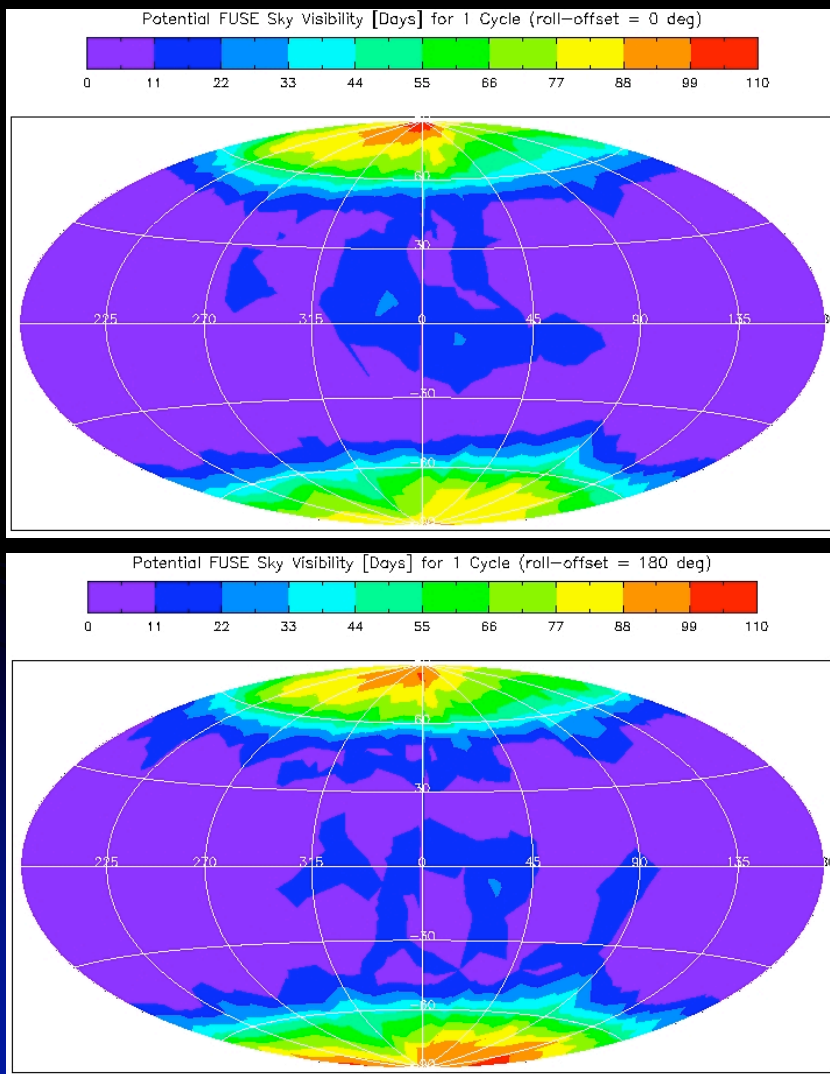
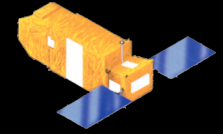
# Future Work/Ideas



- Consider different/better slew algorithms.
  - Different slew algorithms may provide more flexibility and improved sky coverage.
    - Would require further ACS s/w development.
- Continue to improve MP tools.
  - Improved momentum mgmt into Long Range Planning.
  - “SOVA” s/w: Bringing AI techniques to short term scheduling.
  - Longer term, this may be the primary means for performance improvements.
- “180 degree roll” Ops concept.
  - Would place skew axis in radically different direction wrt the B-field at any given time.
  - Recent performance >> thermal conditions may be manageable.
  - Possible show stopper: FES scattered light? (Only at certain beta angles?)
  - Would require further ACS and MP s/w development.

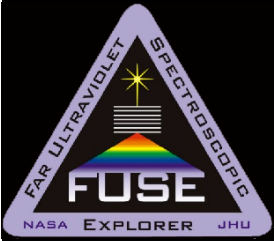


# 180 Roll Sky Coverage

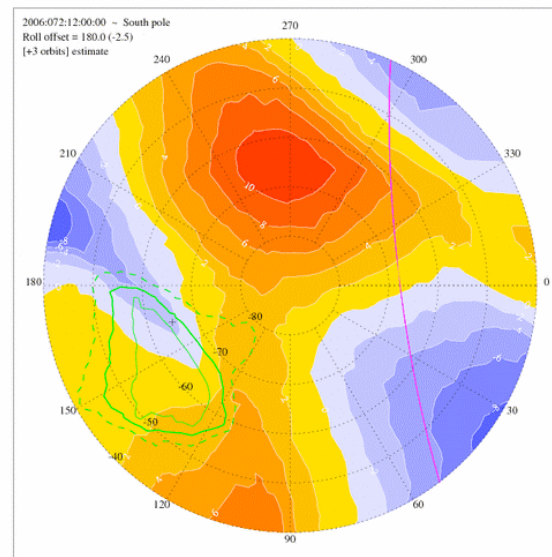
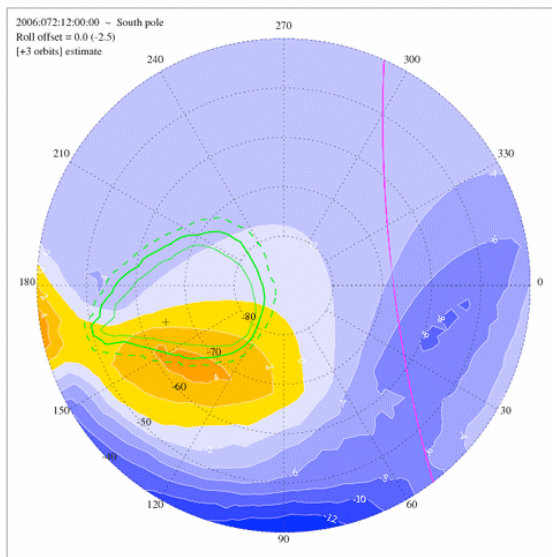
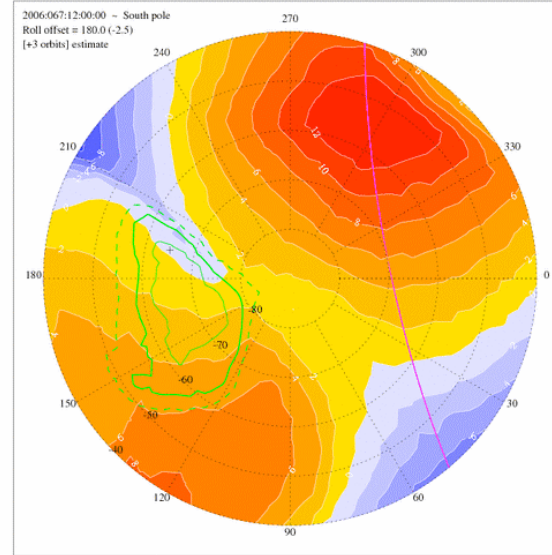
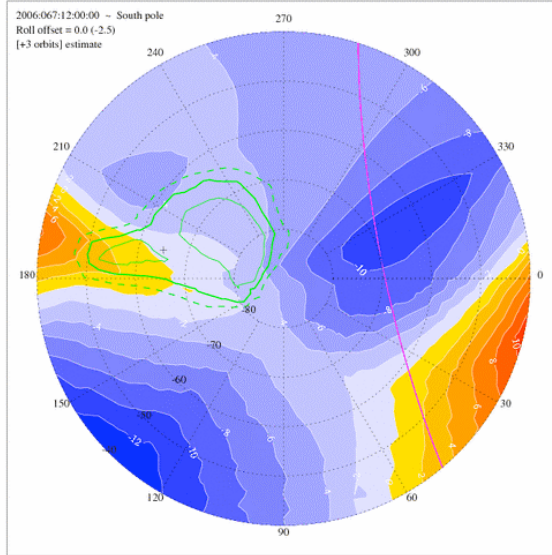
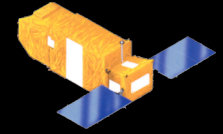


- A 180 roll flip does not do anything dramatic to open up the sky coverage.
  - South cap coverage improves.
- BUT, this is deceptive. It increases the amount of access a given part of the sky, this could still add a lot of flexibility into scheduling.
  - Example: conflicting needs of south (Mag Clouds) and north (long integration targets) might be better accommodated.

(H. Calvani)

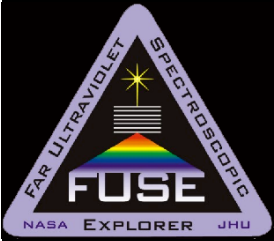


# 180 Roll Momentum/TACO

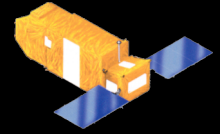


- Momentum pattern changes are dramatic with the roll flip.
- If usable for observing, may increase time we can spend on a given sky region, and may help with long obs.
- (If not usable for observing, may still help tremendously with momentum mgmt activities!)

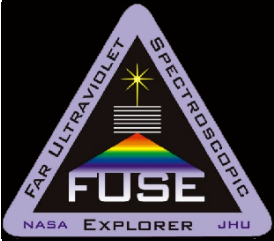
(T. Civeit)



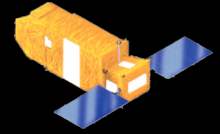
# Issues/Questions-I



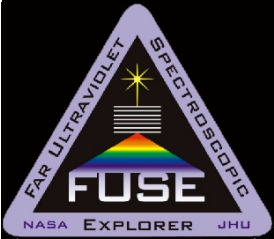
- Individual very high priority targets are expensive drivers of scheduling. (Ex: Attempt to observe Alpha Cen in early March.)
  - Such targets can cost several days to (worst case) a week.
  - Can use effectively all available human resources to plan.
  - Are not guaranteed to succeed! (Often, the real complexities only show up well into the short term scheduling of such observations.)
  - Clear communication needed on designation of priority targets.
- How best to accommodate/prioritize Long Observations.
  - Almost by definition, long observations must be given high priority in order to get scheduled.
  - A single target priority may not be enough. Need a “sliding priority” scheme? (once a target is “started” it becomes higher priority?)
  - May need to access individual targets and lock down best windows manually in the LRP. (But then they become “drivers.” See above.)



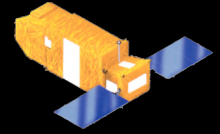
# Issues/Questions-II



- Programs to Address Density of Targets in Accessible Regions of Sky
  - S605/S705 Pole Background rings (pre-defined positions, N&S).
    - Implemented Late-July 2005
  - “U” programs: ~600 previous FUSE targets made available for standardized, plain vanilla LWRS, 8 ks, observation requests.
    - Implemented Mid-August 2005
    - Has reduced the amount of time we spend on backgrounds.
  - Should these be continued? Enhanced? Scaled back?
  - How do we make availability of these data more obvious to the community at large (within our resources!)?

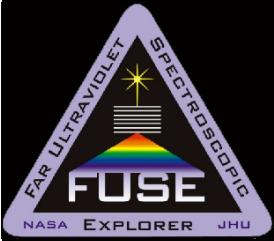


# CaFUSE Status (Van Dixon)

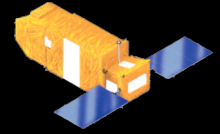


- Draft of CaFUSE article for PASP is circulating. Comments (soon) are welcome.
- Production version of CaFUSE is now v3.1.7; includes numerous minor bug fixes.
- By May 1, plan to release v3.2 together with new effective-area, background, and airglow calibration files.
- v3.2 is expected to be the baseline version, with only calibration file updates in the future.

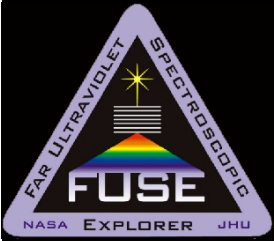




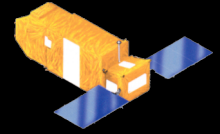
# CaFUSE Version Overview



- v3.0 (Sept. 2004) -- Introduce intermediate-data file.
- v3.1 (Sept. 2005) -- New jitter and high-voltage algorithms reject less good data. Improved treatment of HIST data preserves flux, improves wavelength calibration. Tighter pulse-height limits reduce bkgd.
- v3.2 (May 2006) -- Fix bugs in background-scaling routine and FIFO-overflow correction. Improves LiF1 wavelength zero point ( $\sigma = 2.9$  km/s), pothole correction, and spectral centroiding algorithm. New effective-area files.

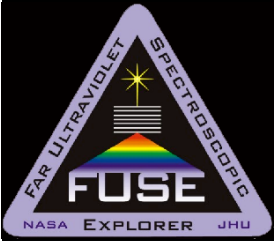


# Data Processing Status

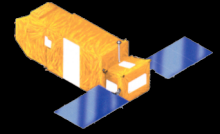


- 4200 observations are now in MAST
- 700 processed with v3.1
- 2800 processed with v3.0
- 700 processed before v3.0

*Data obtained during most of 2000 and the second half of 2001 are being processed without HSKP and JITR files due to SCC hardware and backup tape problems.*



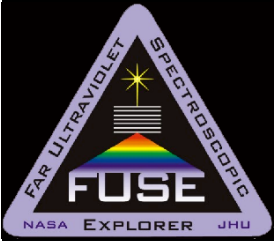
# Data Processing Calendar



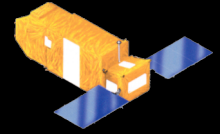
- May 1 -- All data with v3.0 or above
- Oct. 1 -- All data with v3.1 or above
- Jan. 1 -- All data with v3.2

*To meet this schedule, we will devote a third computer to the reprocessing effort.*

*Efforts to recover the 2000 and 2001 HSKP and JITR files will continue.*



# MAST Archive



- Mary Romelfanger has identified about 300 data sets that are not in MAST, but probably should be. Most are in-orbit checkout or calibration observations. She and Dave Sahnou are working on this now.
- *Next step:* Organize web documentation of *FUSE* mission for MAST.

