

## Seven Eight Days in the Life of AR Lac

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### Abstract

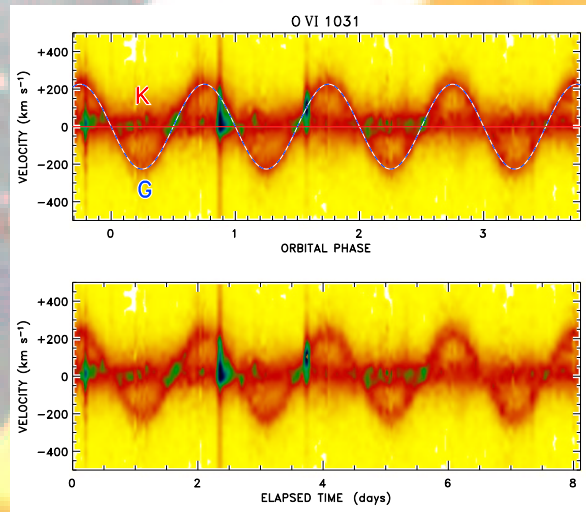
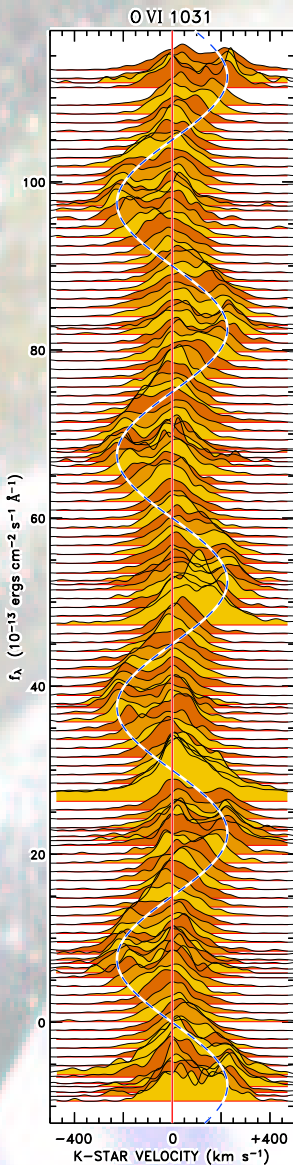
A week long *FUSE* pointing on short period eclipsing binary AR Lacertae (K0 IV+G4 IV;  $P = 2$  d) was coordinated with groundbased Doppler imaging to provide an unprecedented view of structure, energetics, and dynamics of surface activity on a classical RS CVn system.

### Background

AR Lac is an eclipsing binary of K0 and G4 subgiants, separated (surface to surface) by just under the diameter of the larger cooler primary, in a 1.98 d orbit. The K star is slightly more massive than the warmer less evolved secondary. Primary eclipse (K star in front) is total and lasts for several hours; secondary eclipse is partial. At  $d = 42$  pc, AR Lac is one of the brightest coronal stars in the *ROSAT* all sky survey, and a prominent FUV source as well.

### Observations

The *FUSE* pointing was carried out 29 November to 7 December 2004, covering four revolutions of the binary. Total exposure was 225 ks. Emissions of O VI  $\lambda 1031$  and C III  $\lambda 977$  are prominent, and detected from both components of the system (more weakly from G star). Figure *left* illustrates the O VI time series (LiF1a channel), in K-star reference frame (dashed curve traces G-star velocity). Supporting Doppler Imaging was carried out at the Nordic Optical Telescope, McDonald Observatory, and Apache Point. (The multi-site campaign was necessary owing to the nearly exactly 2 day period of the system.)



### Discussion

Figure *above* depicts O VI time series as an image, again in K star frame. Three large O VI flares are conspicuous, with significantly elevated continuum emission. Numerous smaller events also are apparent. Much—but not all—of the flare activity is on the K subgiant primary, and largest outburst ( $\phi \sim 0.86$ ) exhibits strong redshifts. Future analysis will concentrate on flare dynamics, association of events with optical starspots, and behavior during eclipses.



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